

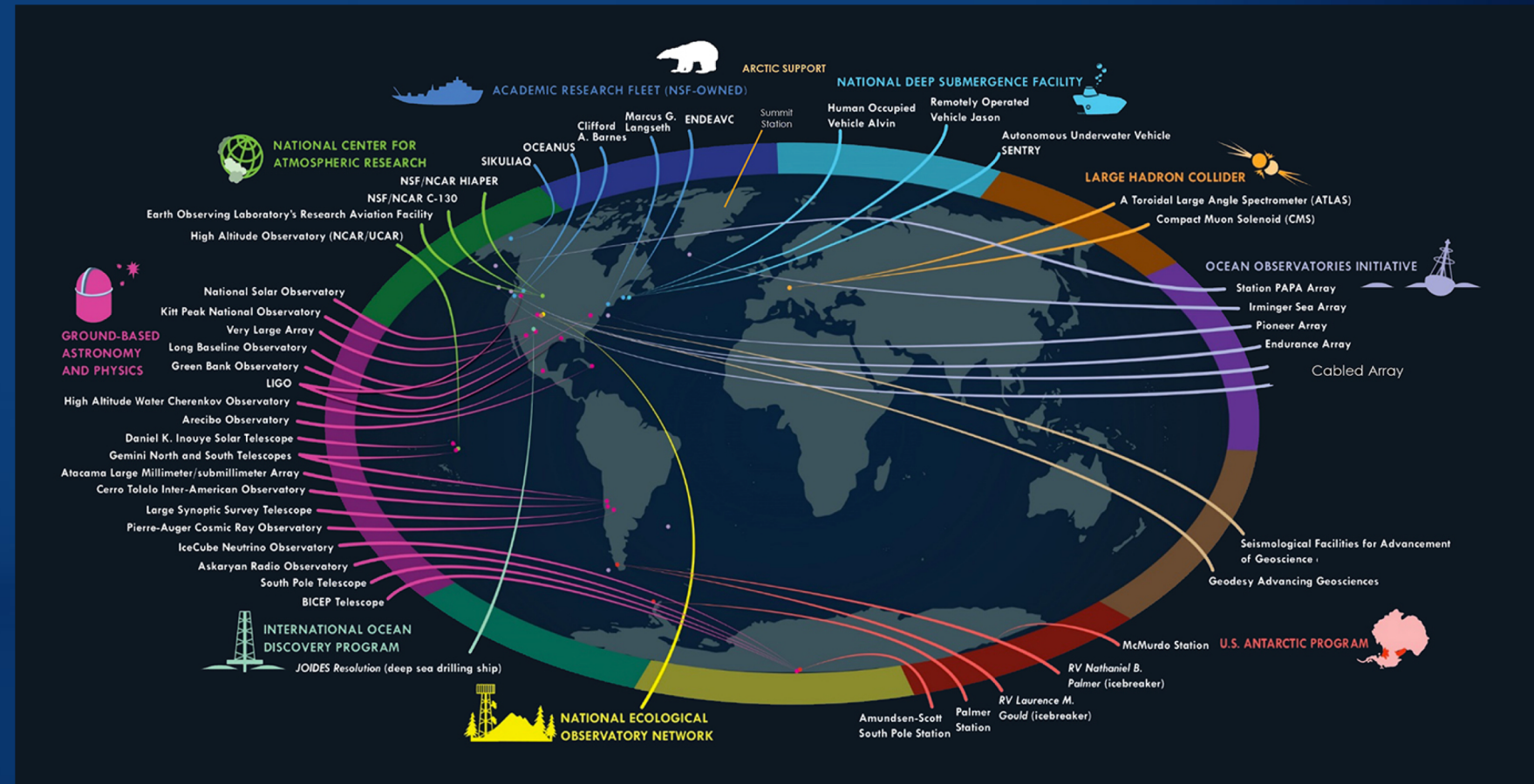
National and Regional Regulatory Structures and how they feed into the International Structure:

USA



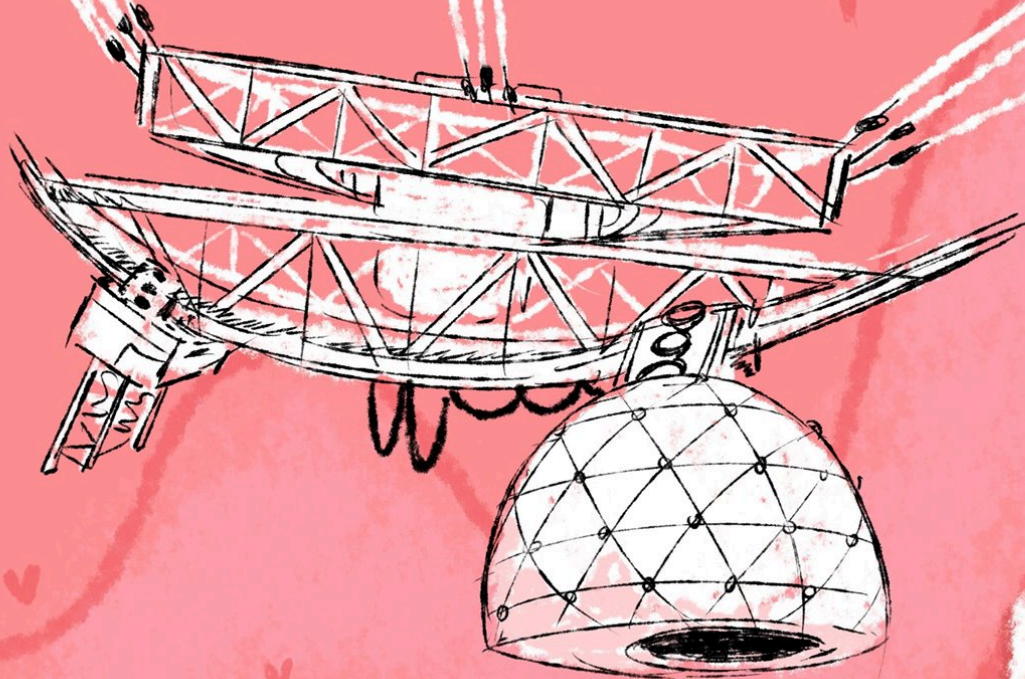
Bevin Ashley Zauderer
Division of Astronomical Sciences
National Science Foundation

March 2, 2020
IUCAF 5th School
Stellenbosch, South Africa



<https://www.nsf.gov/bfa/dcca/csb/>

TO: IUCAF 5th School
FROM: Ashley



ARE YOU RFI?
'CAUSE WE'RE IN
THE SAME
FREQUENCY.

- I. Radio Astronomy: intrinsic value and spectrum needs
- II. Process is complicated: many stakeholders
- III. Getting involved is simple: contribution driven



I. Radio Astronomy: intrinsic value and spectrum needs

II. Process is complicated: many stakeholders

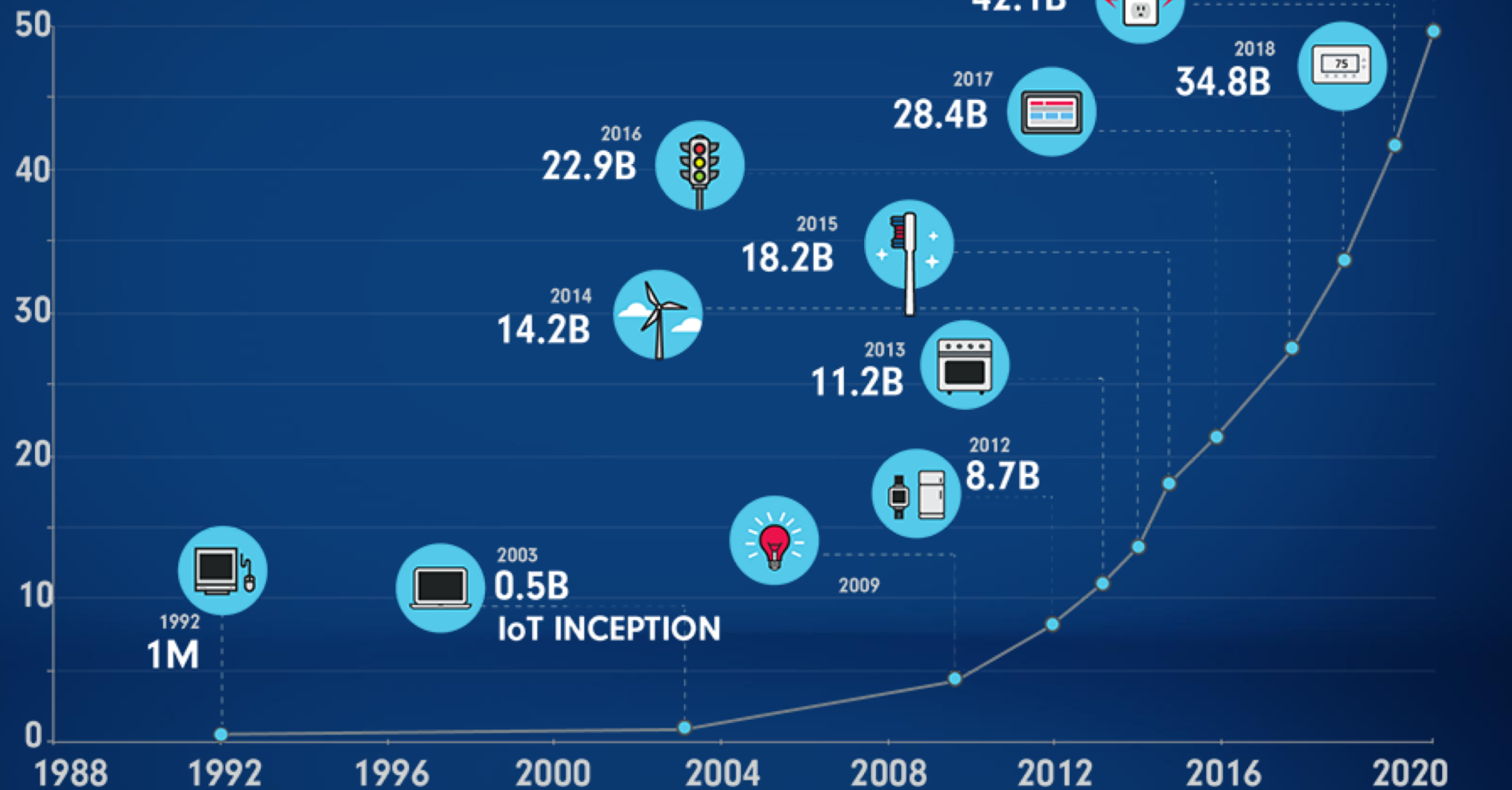
III. Getting involved is simple: contribution driven

“Not everything that *counts* can be counted, and not everything that can be counted *counts*.”

- Einstein



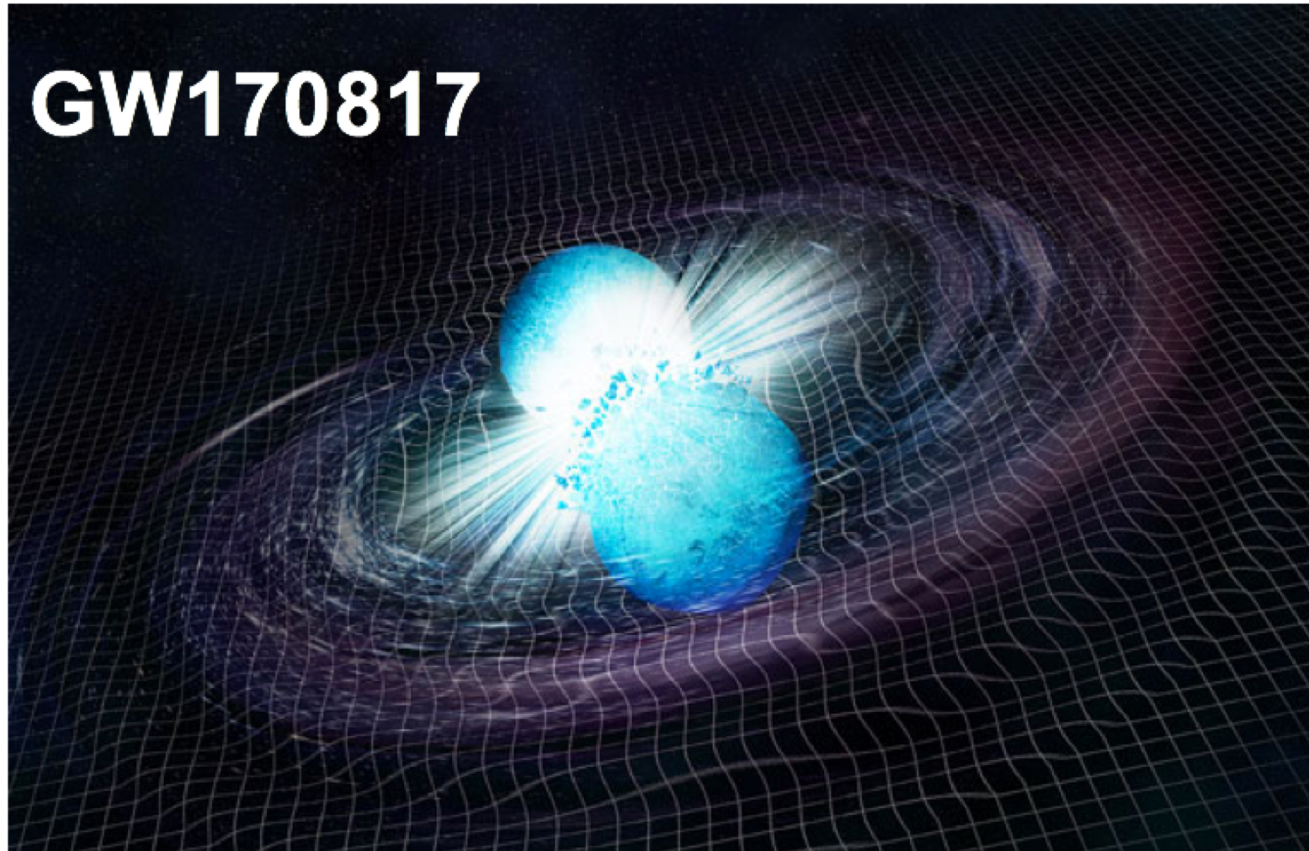
BILLIONS OF DEVICES



Source: Cisco



Multi-messenger & Time-Domain Astronomy



Artist's illustration of the merger of two neutron stars. A new study suggests that the neutron-star merger detected in August 2017 might have produced a black hole.
NASA/CXC/M. Weiss

LIGO – A GIGANTIC INTERFEROMETER

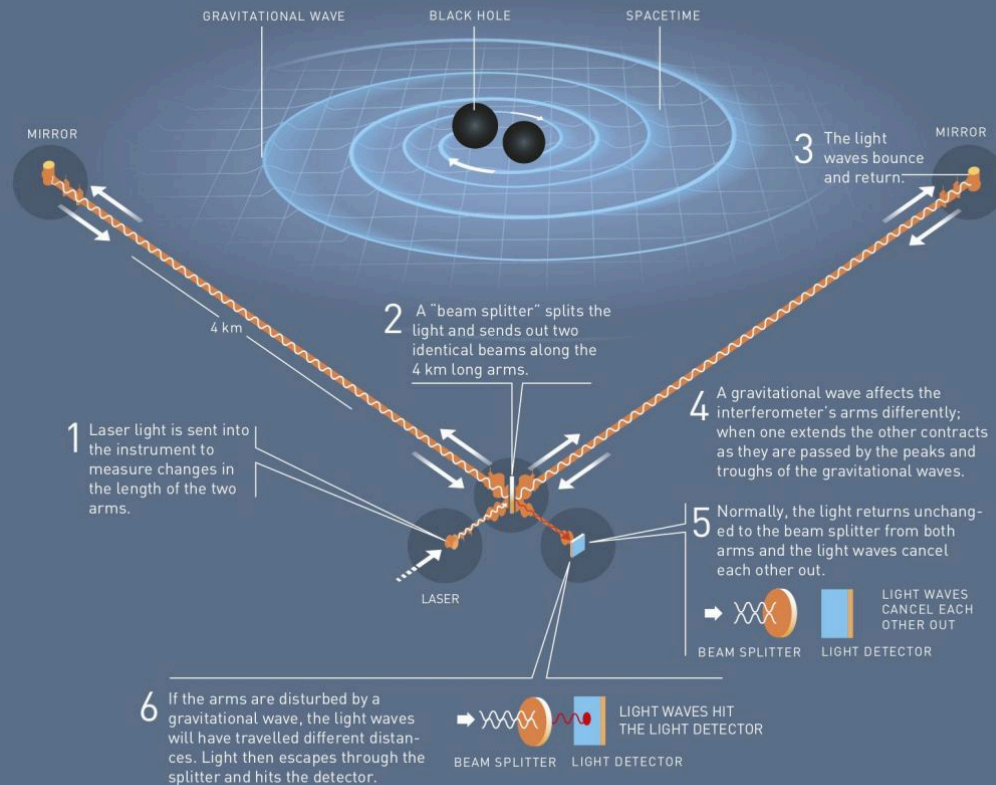


Illustration: © Johan Järnstad/The Royal Swedish Academy of Sciences

A radio counterpart to a neutron star merger

G. Hallinan^{1,*†}, A. Corsi^{2,†}, K. P. Mooley³, K. Hotokezaka^{4,5}, E. Nakar⁶, M. M. Kasliwal¹, D. L. Kaplan⁷, D. A. Frail⁸, S. T. Myers⁸, T. ...

+ See all authors and affiliations

Science 22 Dec 2017:
Vol. 358, Issue 6370, pp. 1579-1583
DOI: 10.1126/science.aap9855

Article

Figures & Data

Info & Metrics

eLetters

 PDF

GROWTH observations of GW170817

The gravitational wave event GW170817 was caused by the merger of two neutron stars (see the Introduction by Smith). In three papers, teams associated with the GROWTH (Global Relay of Observatories Watching Transients Happen) project present their observations of the event at wavelengths from x-rays to radio waves. Evans *et al.* used space telescopes to detect GW170817 in the ultraviolet and place limits on its x-ray flux, showing that the merger generated a hot explosion known as a blue kilonova. Hallinan *et al.* describe radio emissions generated as the explosion slammed into the surrounding gas within the host galaxy. Kasliwal *et al.* present additional observations in the optical and infrared and formulate a model for the event involving a cocoon of



Why does access to the radio spectrum matter?

10 μ Jy at 3 GHz ~2 weeks

2 GHz BW (~1.4 GHz after RFI excision)

<50 MHz is
RAS primary

VLA Observation September 7, 2017

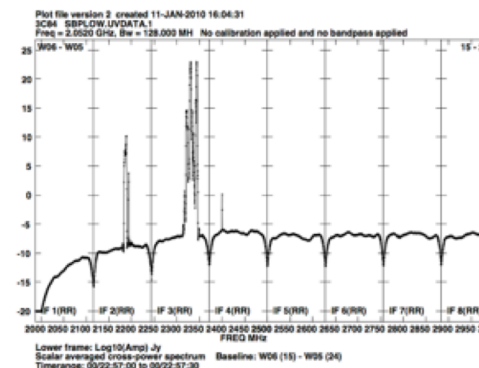
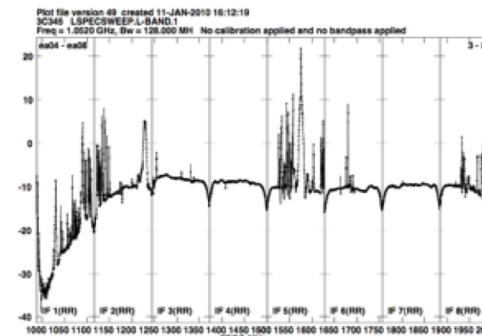


Image Credit: Hallinan et al., Science (2017)



Why does access to the radio spectrum matter?

10 μJy at 3 GHz ~2 weeks

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RAS primary

VLA Observation September 7, 2017

Image Credits: Hallinan et al., Science (16 Oct 2017)

To achieve 2 μJy RMS
requires integration time on source of:

2 GHz bandwidth:
5.5 hours

1.4 GHz bandwidth:
6 hours

50 MHz bandwidth:
185 hours (more than one week)

VLA Exposure Calculator	
Array Configuration	A
Number of Antennas	25
Polarization Setup	<input type="radio"/> Single <input checked="" type="radio"/> Dual
Type of Image Weighting	<input checked="" type="radio"/> Natural <input type="radio"/> Robust
Representative Frequency	3.0000 GHz
Receiver Band	5
Approximate Beam Size	0.977"
Digital Samplers	<input type="radio"/> 3 bit <input checked="" type="radio"/> 8 bit
Elevation	Medium (25-50 degrees)
Average Weather	Autumn
Calculation Type	<input checked="" type="radio"/> Time <input type="radio"/> BW <input type="radio"/> Noise/Tb
Time on Source (UT)	1.1248w
Total Time (UT)	1.4184w
Bandwidth (Frequency)	50.0000 MHz
Bandwidth (Velocity)	4,996.5410 km/s
RMS Noise (units/beam)	2.0000 μJy



Exposure is too long

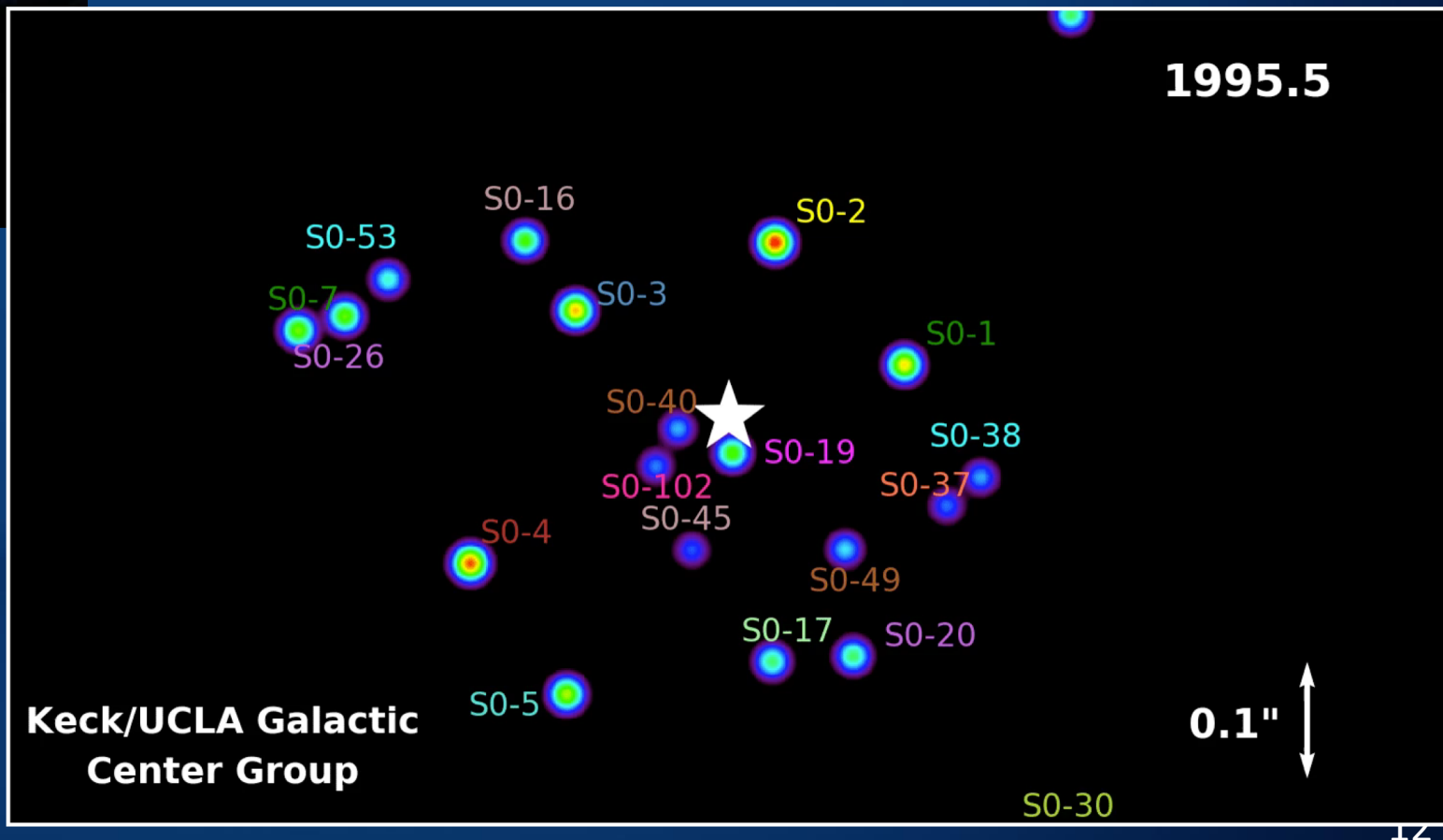
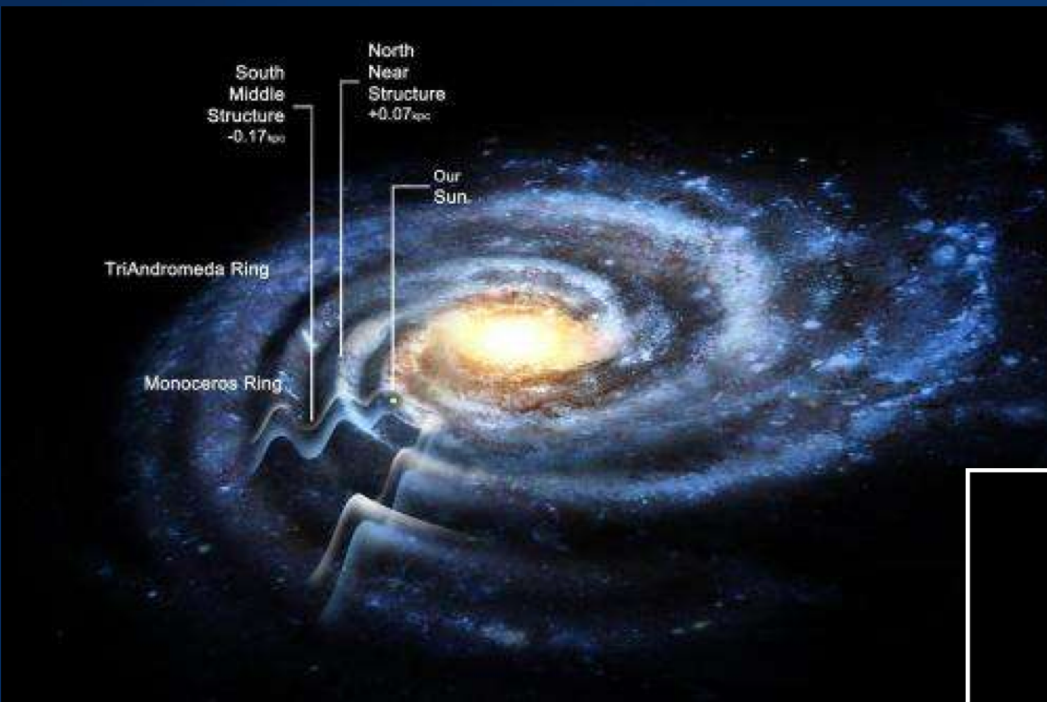
That is a lot of VLA time on one source. You may want to change your values for noise and bandwidth.

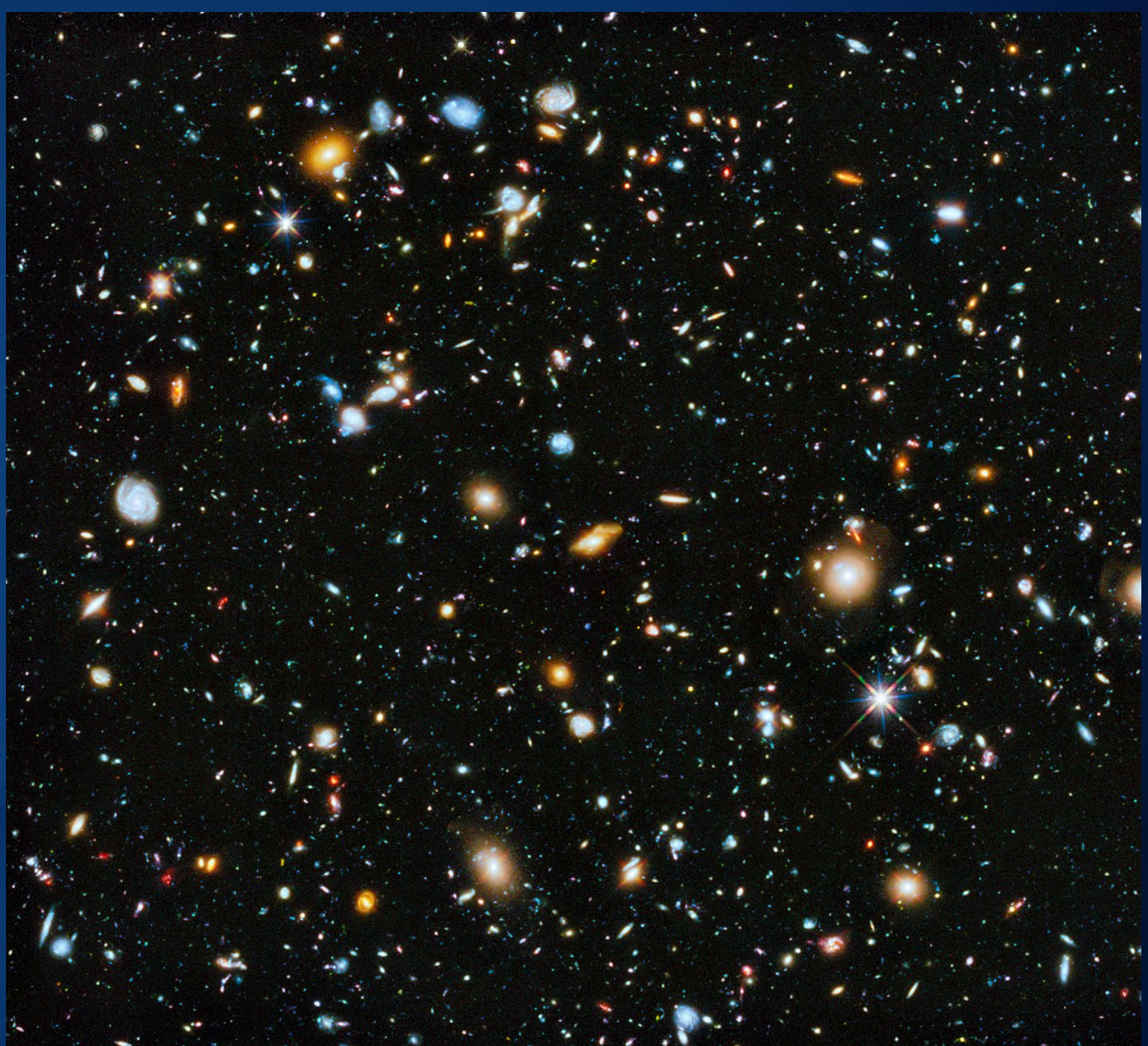
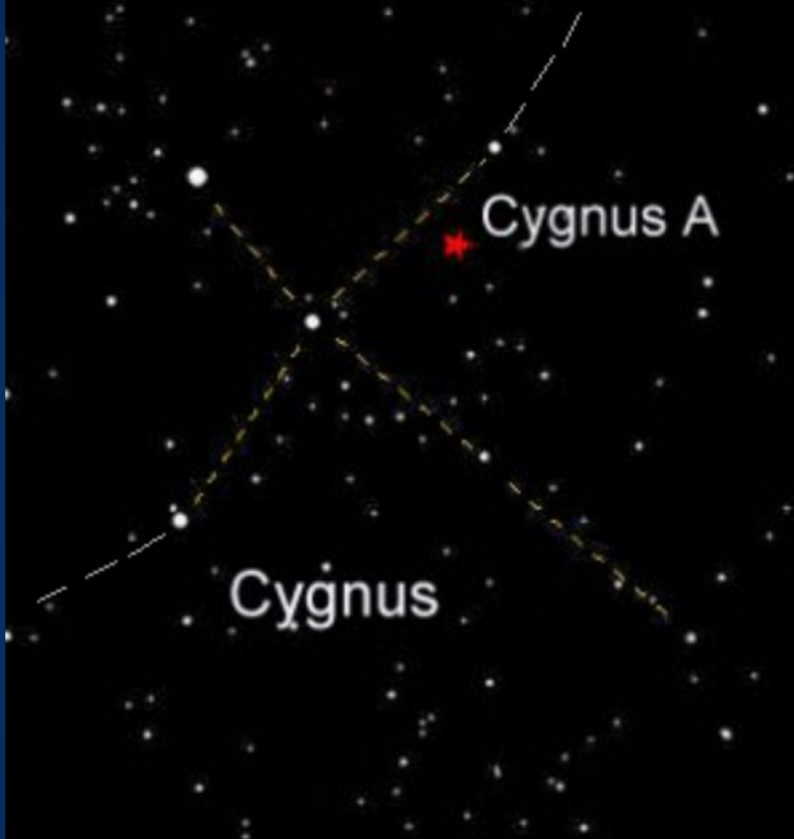


Image credit: Event Horizon Telescope Collaboration

Why care about black holes?







Hubble Ultra Deep Field, NASA

Radio Galaxies – precision astrometry with VLBI

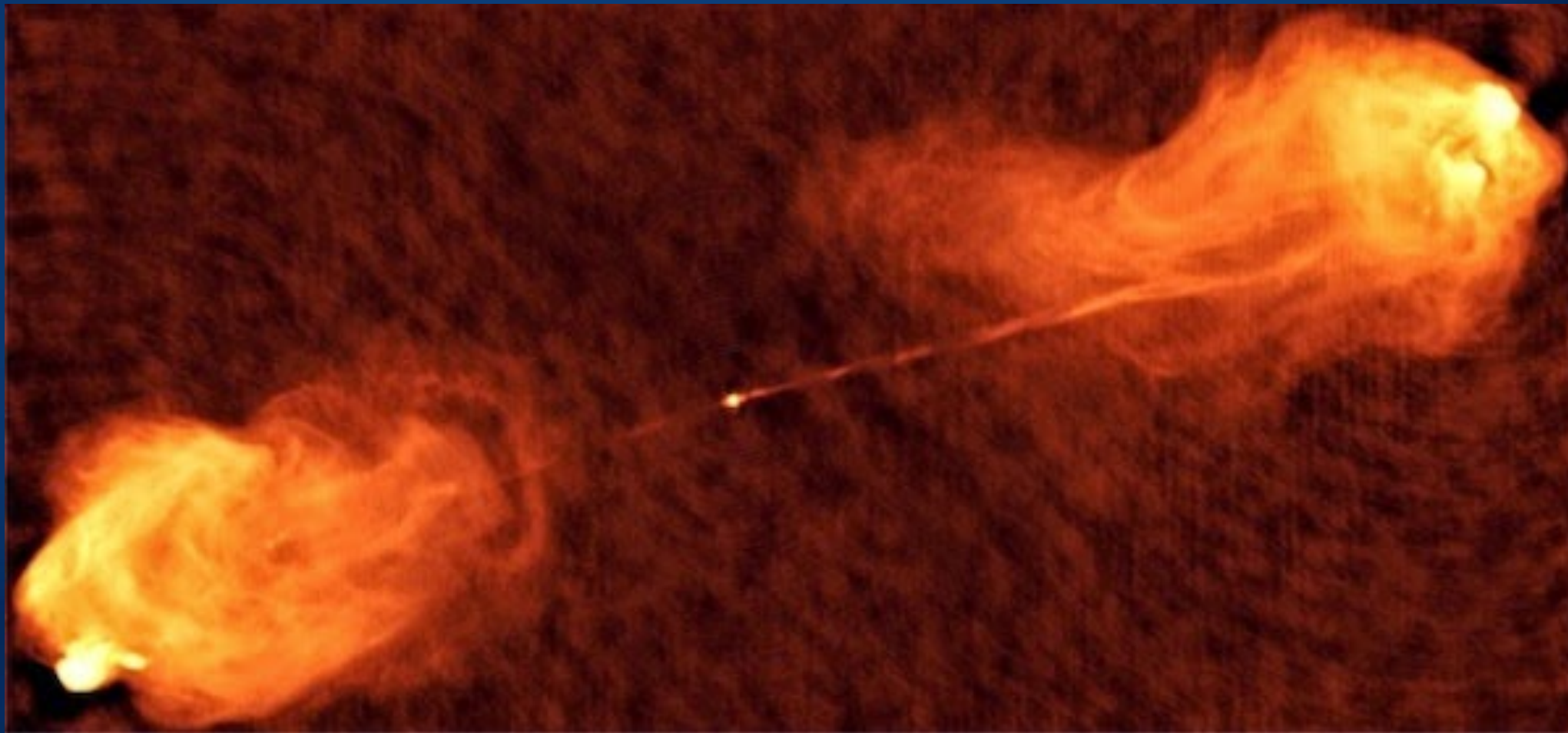
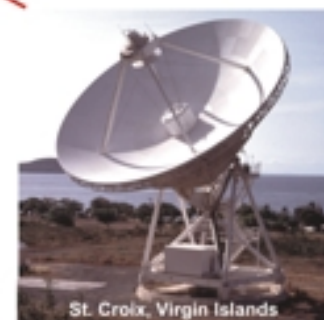
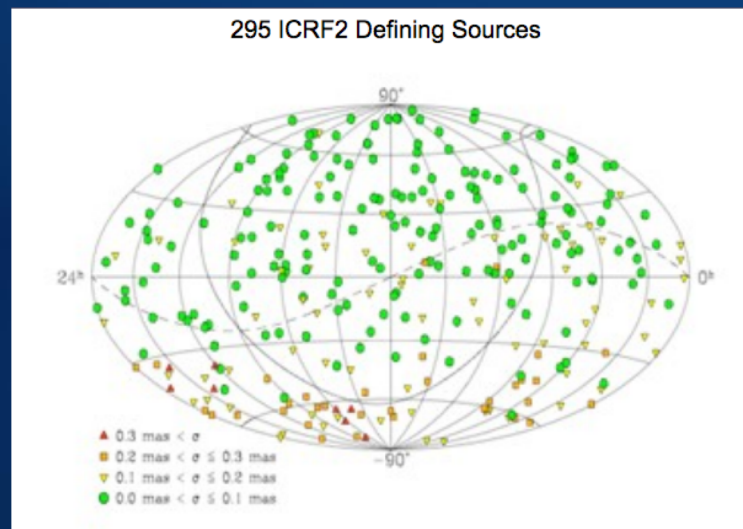
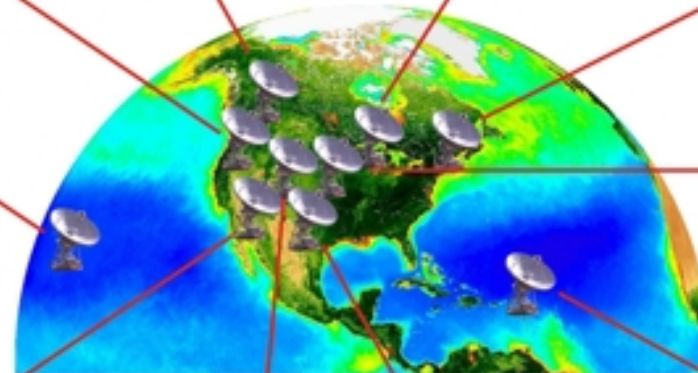
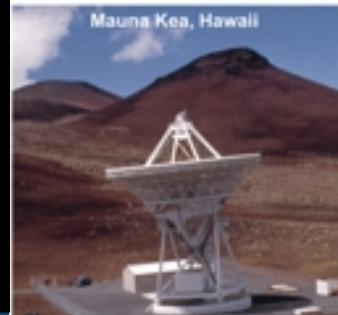


Photo Credit: NRAO/AUI



International Celestial Reference Frame



Very Long Baseline Array

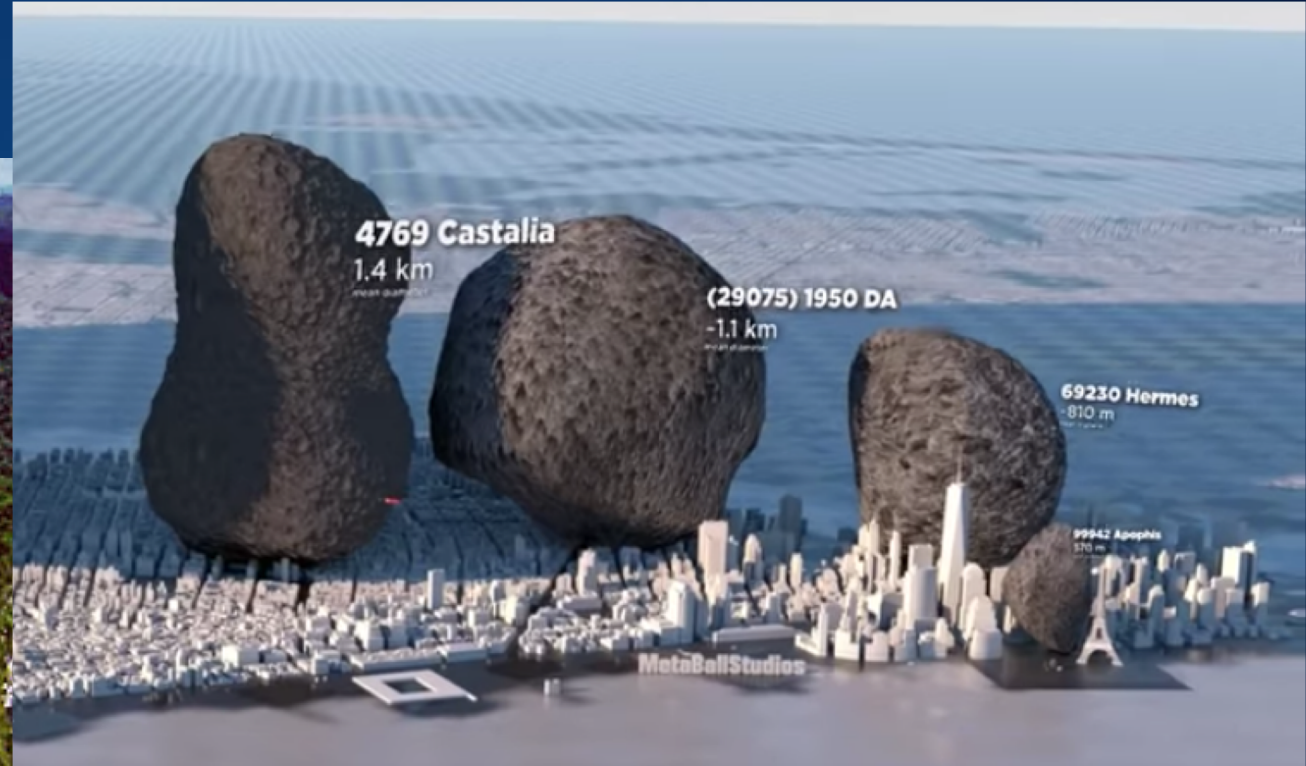
Credits: D. Boboltz, K. Johnston (US Naval Observatory), bestanimations.com, NRAO



Near Earth Objects (NEOs)



Arecibo Observatory, Puerto Rico

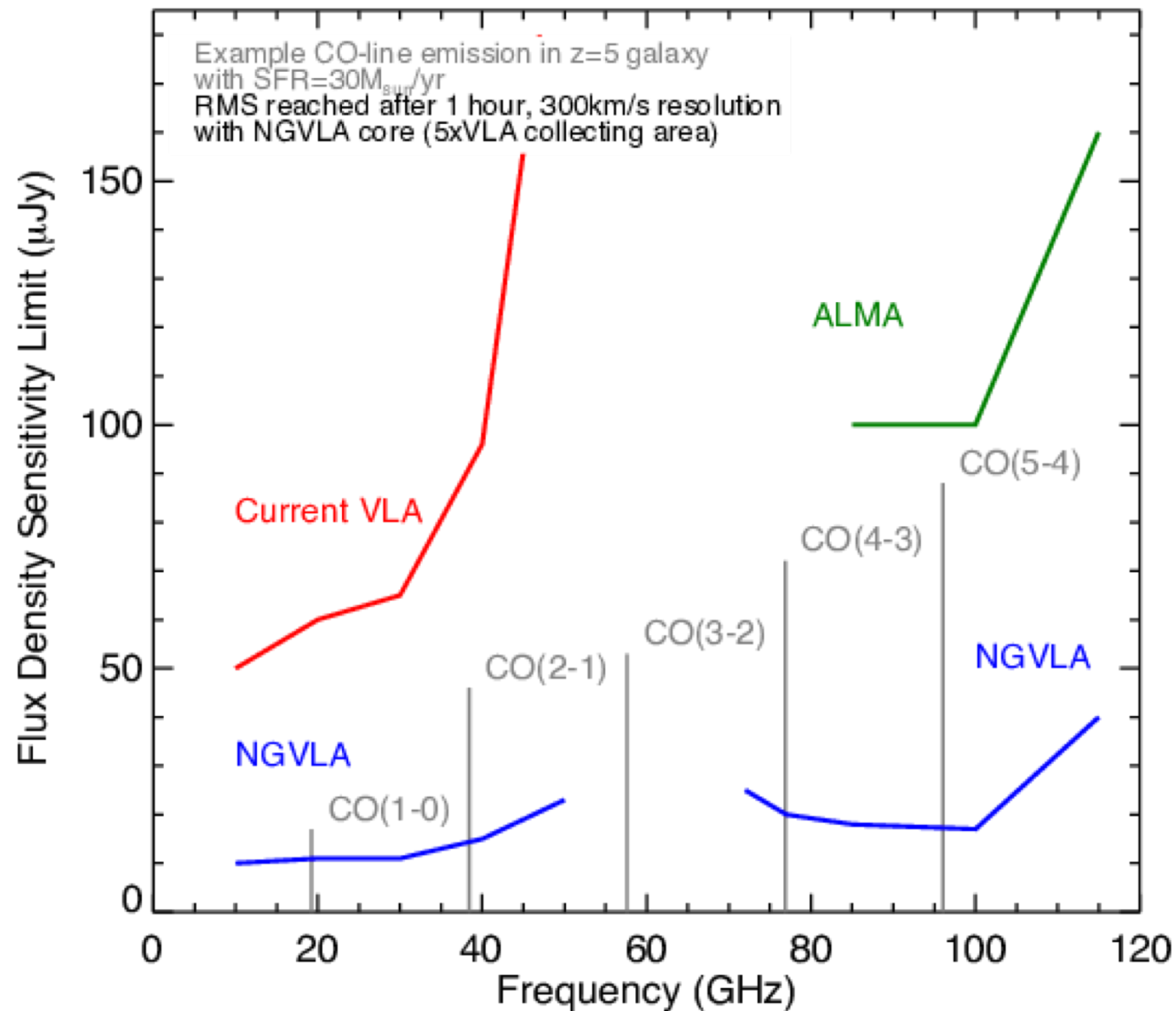


<https://laughingsquid.com/size-of-asteroids-compared-to-new-york-city/>



Green Bank Telescope, Green Bank, WV





Spectrum Frontiers > 24 GHz

<https://www.fcc.gov/document/spectrum-frontiers-ro-and-fnprm>

Spectrum Horizons > 95 GHz

<https://www.fcc.gov/document/fcc-opens-spectrum-horizons-new-services-technologies>



Science with an ngVLA. Cold gas in High- z Galaxies: The molecular gas budget - Decarli, R. *et al.* ASP Conf.Ser. 517 (2018) 565
arXiv:1810.07546 [astro-ph.GA]

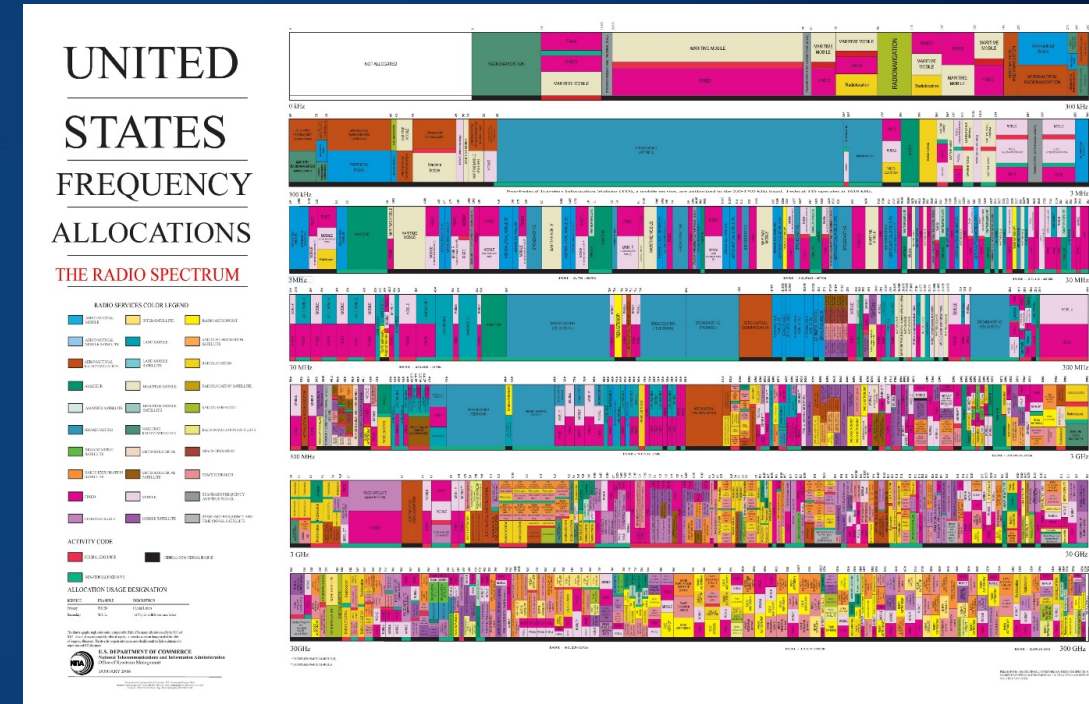
The RFI environment as we know it is changing... *rapidly*.

Science missions are requiring expanded spectrum usage.



Frequency Usage Takeaways

- Protected frequency bands include most important identified spectral lines for studying the local universe (e.g. HI, CO, OH masers), but doppler-shifted lines from sources further away in the Universe fall into non-protected bands. Frequencies used for observation are often non-interchangeable, and much observation is done opportunistically.

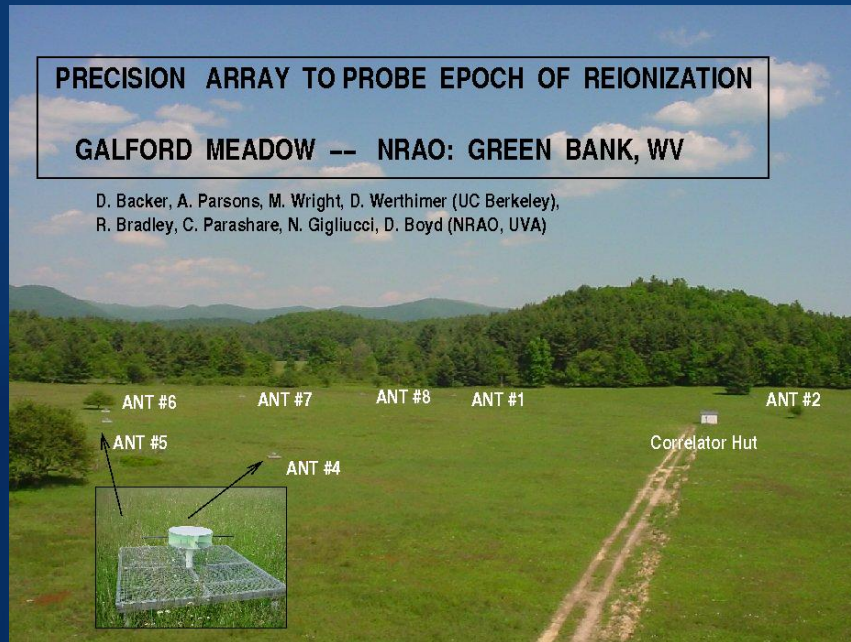


Epoch of Reionization

HI: 21 cm \rightarrow 1.5 m

Freq \sim 1420 MHz \rightarrow 200 MHz

$$1 + z = \frac{f_{\text{emit}}}{f_{\text{obsv}}}$$



PRECISION ARRAY TO PROBE EPOCH OF REIONIZATION

GALFORD MEADOW -- NRAO: GREEN BANK, WV

D. Backer, A. Parsons, M. Wright, D. Werthimer (UC Berkeley),
R. Bradley, C. Parashare, N. Gigliucci, D. Boyd (NRAO, UVA)

Image Credit: w.astro.berkeley.edu

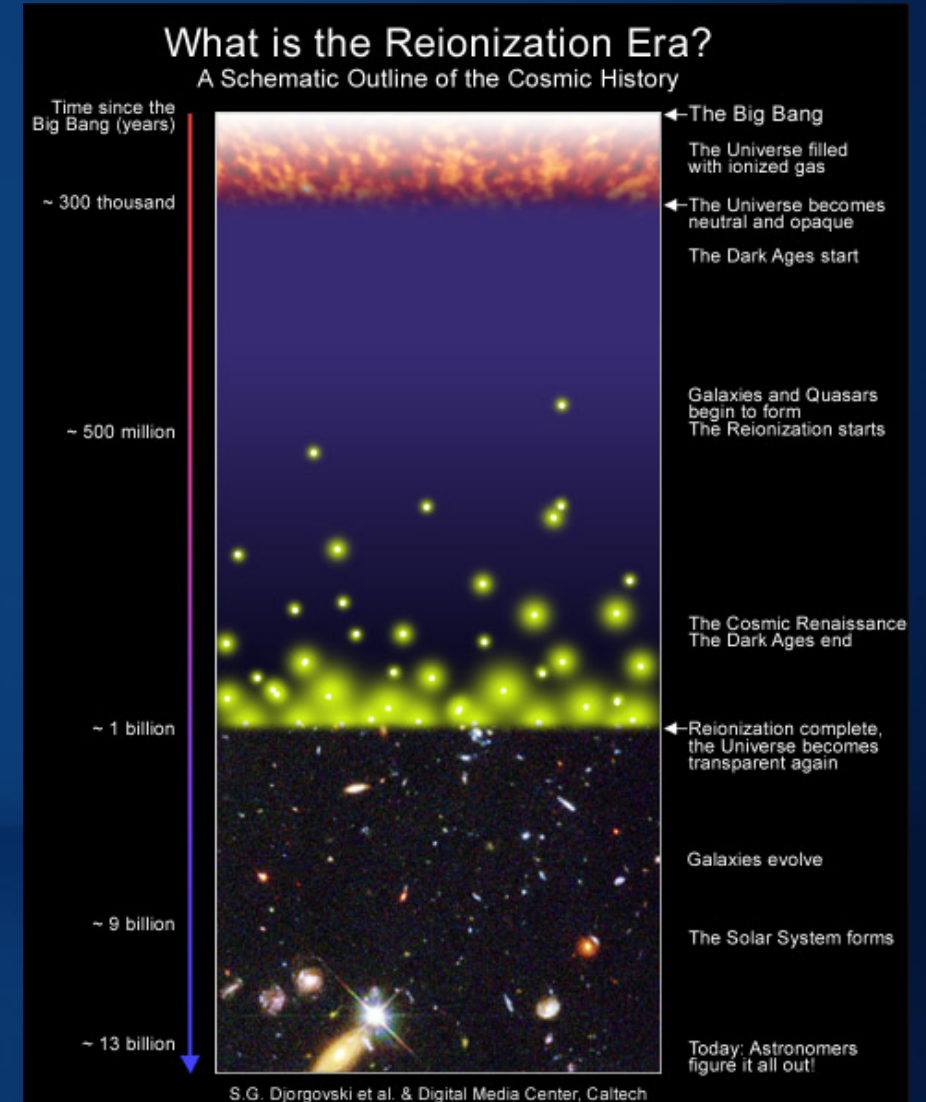
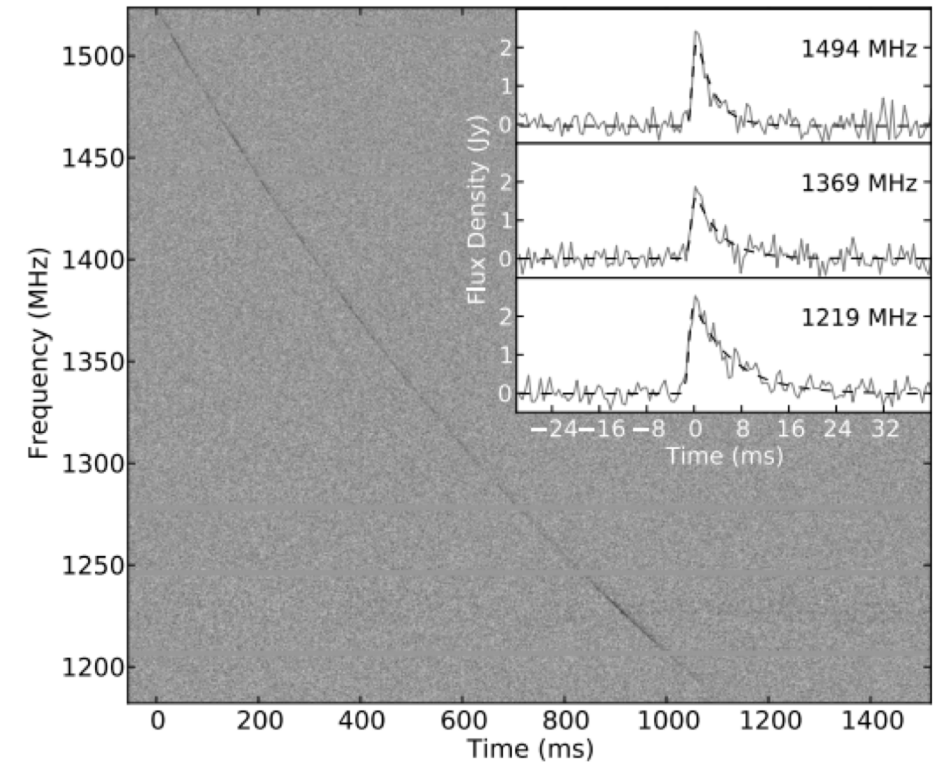


Image Credit: Djorgovski et al. (Caltech); www.haystack.mit.edu



Frequency Usage Takeaways

- It is imperative that the increasing demands for spectrum take into consideration the challenges to scientific progress; efforts to coordinate and to limit out-of-band emissions are crucial; Astronomy observations also include continuum emission (thermal, non-thermal).



Thornton et al., 2013, Science



Table 1: Overall EVLA Performance Goals

Parameter	VLA	EVLA	Factor
Continuum Sensitivity (1- σ , 9 hr)	10 μ Jy	1 μ Jy	10
Maximum BW in each polarization	0.1 GHz	8 GHz	80
Log (Frequency Coverage over 1–50 GHz)	22%	100%	5



Table and Image
Credit: NRAO



Frequency Usage Takeaways



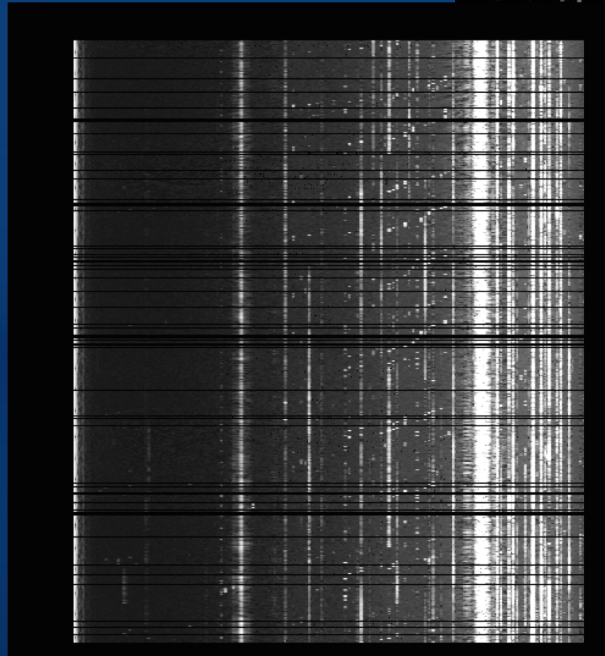
Image credit: almaobservatory.org

- The United States has significant scientific assets / large facilities outside of its national borders.
- Observatories tend to be in geographically remote sites, but radio emission from moving emitters: car radars, satellites and high altitude delivery systems will be an increasing challenge.



What is coming...

- Constellations of thousands of satellites (20-50 GHz regime) such that from any location you would always “see” at least one, preferably (in mind of satellite providers) up to 3 or 4 satellites
- Mobile telecommunications
- High Altitude Platform Systems



How do we preserve access to the spectrum for the next decade and beyond?

- **NRQZ (established 1958) needs updated protections from airborne transmitters; other radio telescopes need also need newly established quiet/coordination zones**



The National Radio Quiet Zone (NRQZ) was established by the Federal Communications Commission (FCC) in [Docket No. 11745](#) (November 19, 1958) and by the Interdepartment Radio Advisory Committee (IRAC) in Document 3867/2 (March 26, 1958) to minimize possible harmful interference to the National Radio Astronomy Observatory (NRAO) in Green Bank, WV and the radio receiving facilities for the United States Navy in Sugar Grove, WV. The NRQZ is bounded by NAD-83 meridians of longitude at 78d 29m 59.0s W and 80d 29m 59.2s W and latitudes of 37d 30m 0.4s N and 39d 15m 0.4s N, and encloses a land area of approximately 13,000 square miles near the state border between Virginia and West Virginia.

Credit: Green Bank Observatory



Credit: NRAO



- I. Radio Astronomy's Value and Spectrum Needs
- II. Process is complicated: many stakeholders
- III. Getting involved is simple: contribution driven



US Domestic Spectrum Policy

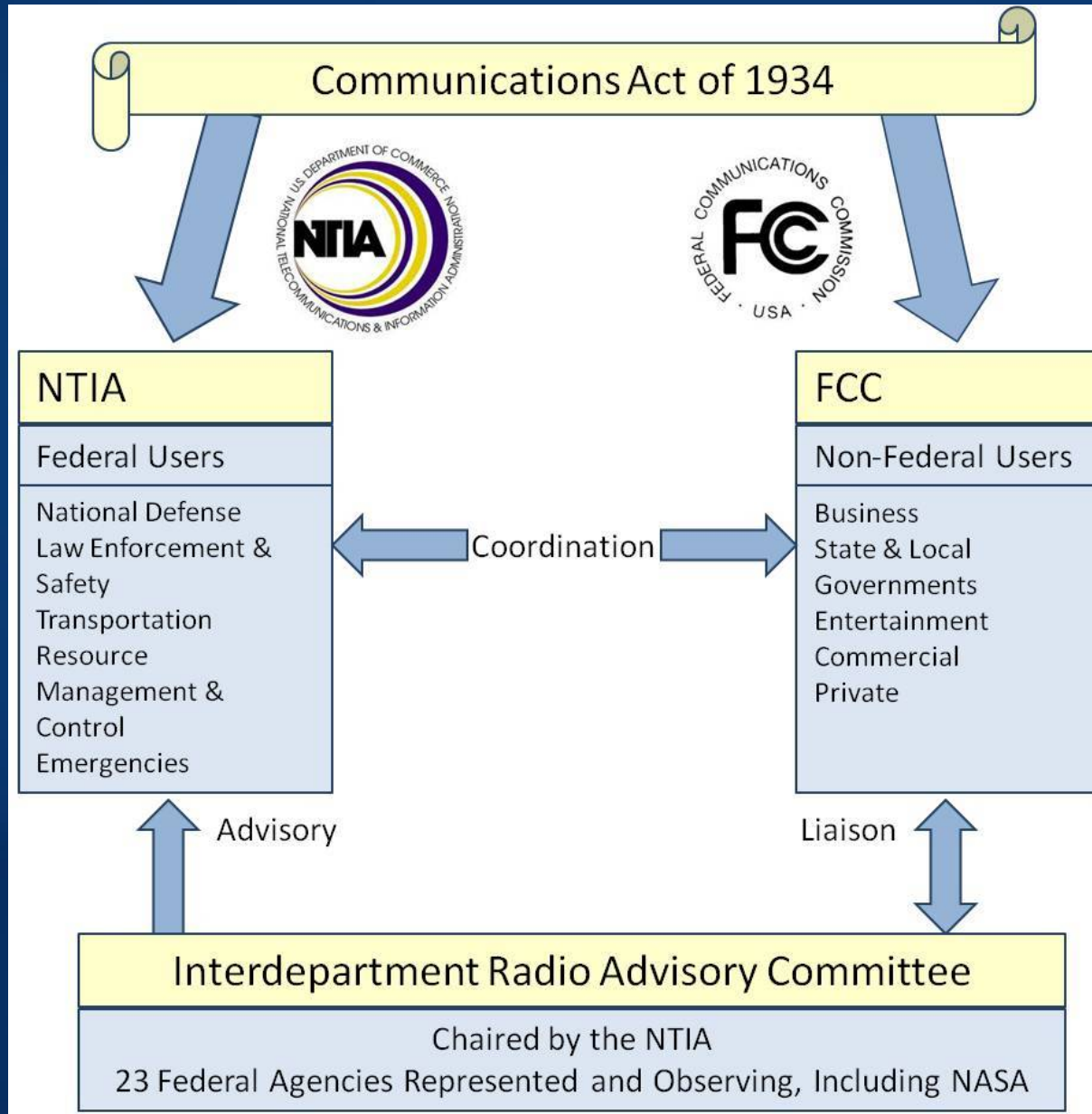
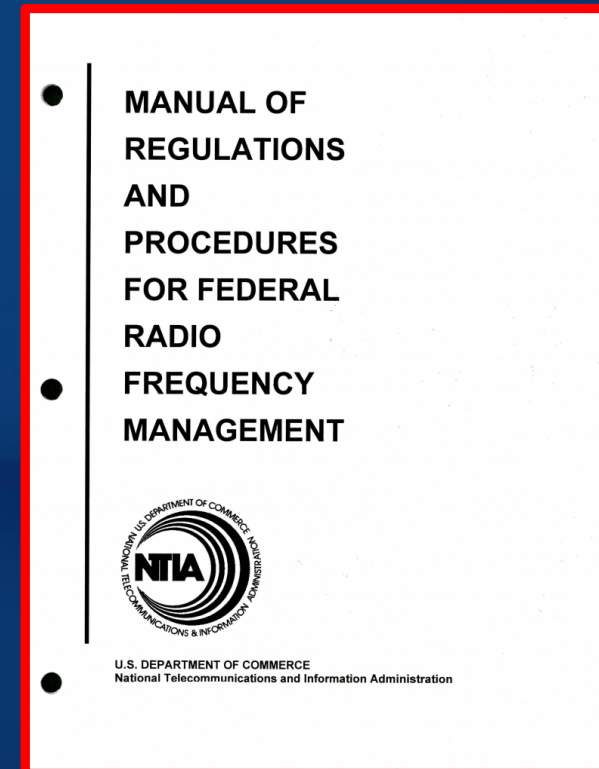


Image Credit:
www.nasa.gov
National Spectrum
Management Plan



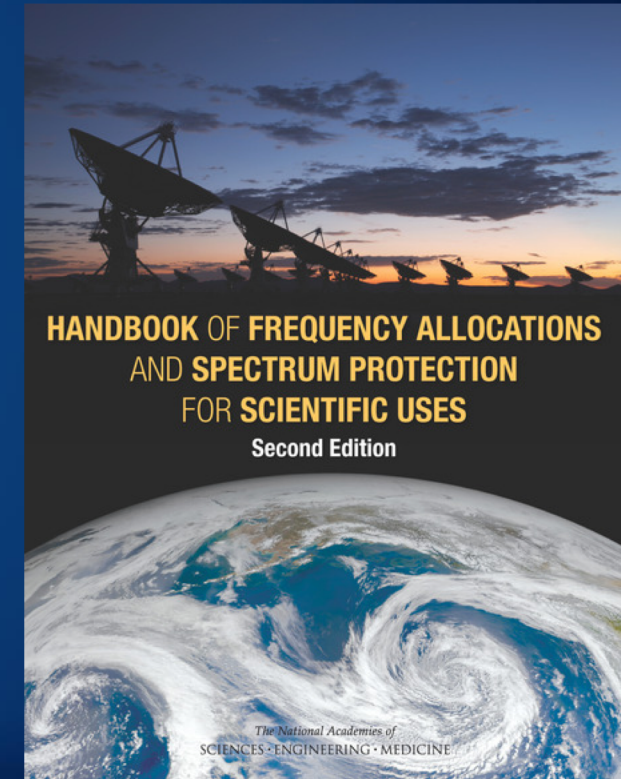
Allocations and Coordination

- Radio Regulations:
 - (1) International (ITU-R Radio Regulations; www.itu.int)
 - (2) National (USA: NTIA - www.ntia.doc.gov; FCC - www.fcc.gov)



Astronomical Input to formulating NSF's position

- **CORF** – Committee on Radio Frequencies, National Academies
 - Chair: Liese van Zee
- **IUCAF** – International Committee working in field of spectrum management on behalf of passive radio sciences (set up in 1960 by URSI, IAU and COSPAR; <http://www.iucsf.org/>)
 - IUCAF Secretariat: Harvey Liszt
- **CRAF** – European Science Foundation's Committee on Radio Astronomy Frequencies, 22 member countries (<https://www.craf.eu>)
 - Chairman: Michael Lindquist; Frequency Manager: Waleed Madkour
- **Spectrum Managers** at telescopes and/or their Managing Organizations
 - NRAO: Harvey Liszt
 - Very Large Array: Dan Mertely
 - Arecibo: Angel Vazquez
 - Green Bank: Paulette Woody
- **Astronomers and Telescope Staff**



NSF ESM Unit Activities



Federal
Communications
Commission

- At NSF – Coordination with other Directorates and Divisions with spectrum needs, manage spectrum related grants portfolio (CORF, Enhancing Access to the Radio Spectrum Solicitation; EARS)
- Coordinate with other US Agencies
- Interface with commercial interests to advocate for their taking “practicable” steps to not cause interference to passive services

NSF ESM Unit Activities



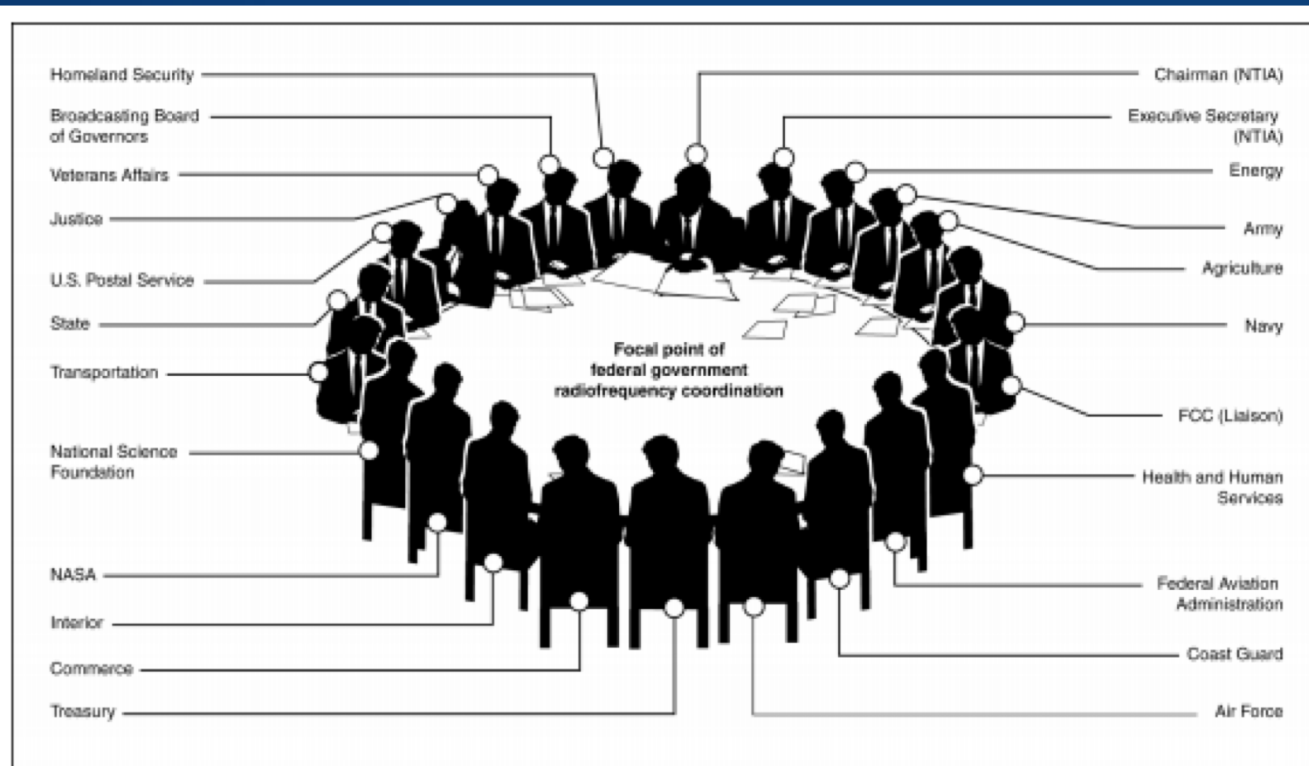
OAS | CITEL



- Represent NSF as a Federal Agency to the National Telecommunications and Information Administration
 - 10 subcommittees including
 - IRAC
 - FAS (NRQZ coordination)
- Representation on official U.S. Delegations to the Inter-American Telecommunications Commission (CITEL) of the Organization of American States (OAS)
- Representation on official U.S. Delegations to the International Telecommunication Union's World Radiocommunication Conference (WRC 2019), including leading 7D – Radio Astronomy

Interdepartment Radio Advisory Committee

- FAS - Frequency Assignment Subcommittee
- SPS - Spectrum Planning Subcommittee
- RCS - Radio Conference Subcommittee
- SSS - Space Systems Subcommittee
- TSC - Technical Subcommittee
- EPS - Emergency Planning Subcommittee
- PPSG – Policy and Plans Steering Group
- Ad Hocs: US-Mexico, US-Canada, NTIA manual modernization



Source: NTIA.

- **MANUAL OF REGULATIONS AND PROCEDURES FOR FEDERAL RADIO FREQUENCY MANAGEMENT**



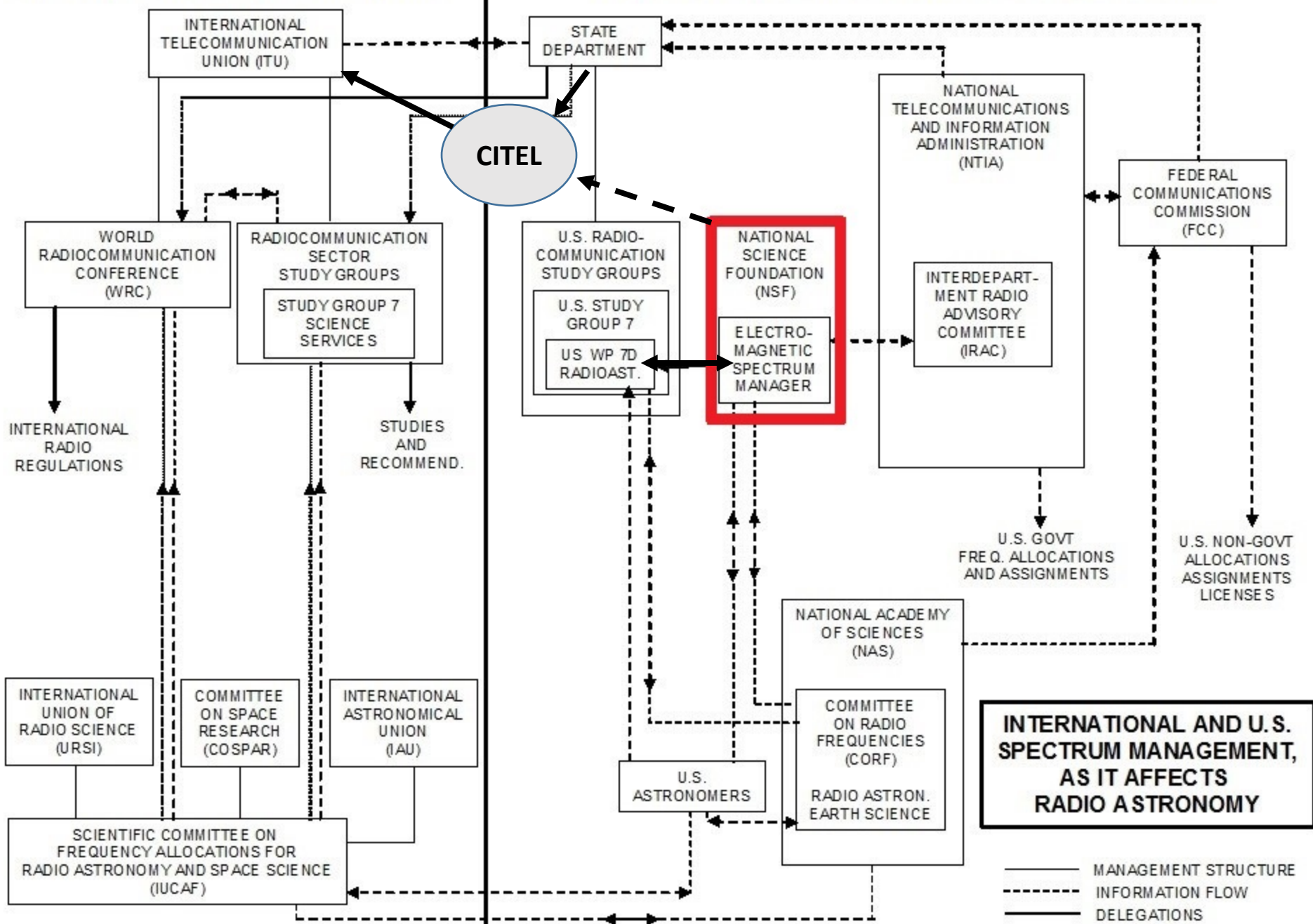
U.S. DEPARTMENT OF COMMERCE
National Telecommunications and Information Administration

•



INTERNATIONAL ORGANIZATIONS

UNITED STATES OF AMERICA ORGANIZATIONS





World Radiocommunication Conference



- 2019 World Radiocommunication Conference (WRC-19) in Sharm El-Sheikh, Egypt
- Technical preparatory work done in the ITU Radiocommunication Sector Study Groups (ITU-R)
- US Regulators oversee conference preparations by Federal Government (NTIA) and private sector (FCC)
- Conference Preparatory Meeting (CPM) report contained approaches (Methods) for satisfying each agenda item (technical basis upon which Administration proposals are made)





Radiocommunication Study Groups

- www.itu.int/en/ITU-R/study-groups
- SG 1: Spectrum Management
- SG 3: Radiowave Propagation
- SG 4: Satellite Services
- SG 5: Terrestrial Services
- SG 6: Broadcasting Services
- SG 7: Science Services
 - Working Party 7A – Time signals and frequency standards
 - Working Party 7B – Space Radiocommunication applications
 - Working Party 7C – Remote sensing systems
 - Working Party 7D – Radio astronomy

Bi-annual meetings in Geneva for all Study Groups and Working Parties, monthly national preparatory meetings leading up to International meetings



- I. Radio Astronomy: intrinsic value and spectrum needs
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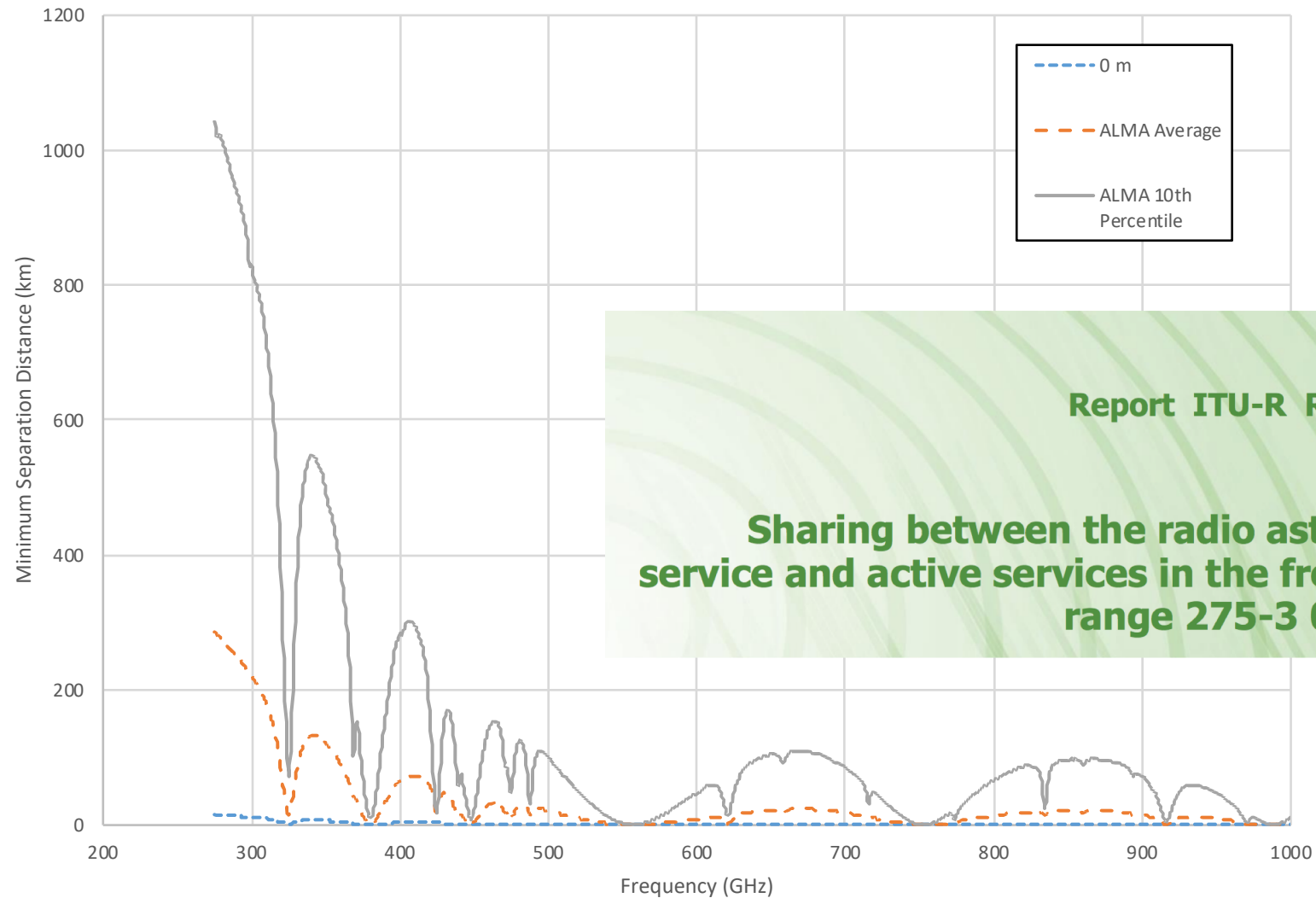
“If [we] fail to plan, [we] are planning to fail!” – B. Franklin

Cooperation and Teamwork is vital



World Radio Conference 2019 Agenda Item 1.15: mm-wave

<http://www.itu.int/pub/R-REP-RA.2189-1-2018>



Report ITU-R RA.2189-1
(09/2018)

**Sharing between the radio astronomy
service and active services in the frequency
range 275-3 000 GHz**

Updated Fall 2018



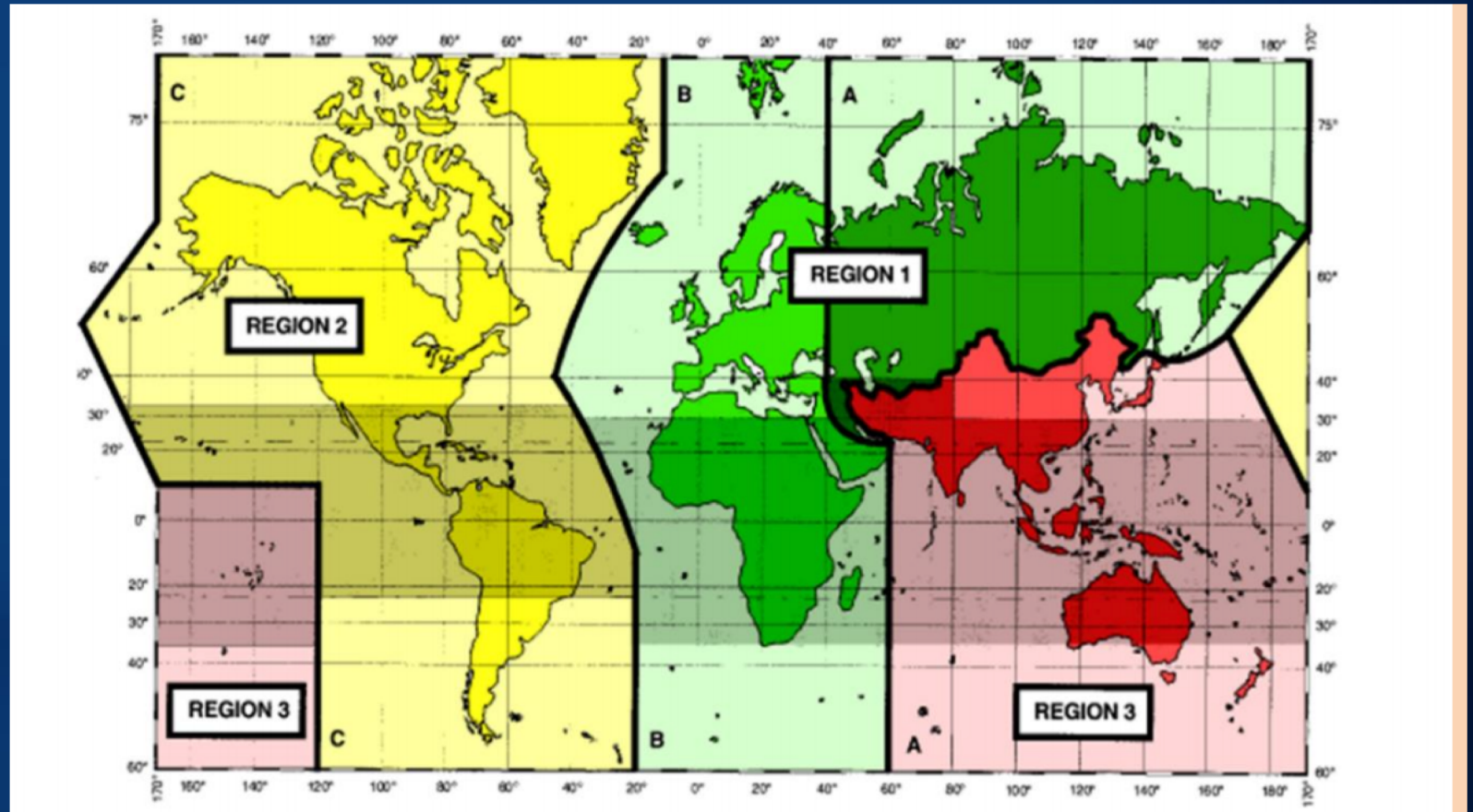
National and Regional Regulatory Structures and how they feed into the International Structure:

Americas, CITE



Bevin Ashley Zauderer
Division of Astronomical Sciences
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March 2, 2020
IUCAF 5th School
Stellenbosch, South Africa



Source: ITU



Member States

All 35 independent states of the Americas have ratified the [OAS Charter](#) and are members of the Organization.

	Antigua and Barbuda		Costa Rica		Haiti		Saint Lucia
	Argentina		Cuba ¹		Honduras		Saint Vincent and the Grenadines
	Barbados		Dominica (Commonwealth of)		Jamaica		Suriname
	Belize		Dominican Republic		Mexico		The Bahamas (Commonwealth of)
	Bolivia		Ecuador		Nicaragua		Trinidad and Tobago
	Brazil		El Salvador		Panama		United States of America
	Canada		Grenada		Paraguay		Uruguay
	Chile		Guatemala		Peru		Venezuela (Bolivarian Republic of)
	Colombia		Guyana		Saint Kitts and Nevis		

Comisión Interamericana de Telecomunicaciones (CITEL)

Mission Statement



Organization of
American States

Vision

The full integration of the American States into the World Information Society and the digital economy, with a view to enabling and accelerating social, economic, cultural, and environmentally sustainable development for all the region's inhabitants through the development of telecommunications and information and communication technologies (ICTs).

Mission

To facilitate and promote the integral and sustainable development of interoperable, innovative, and reliable telecommunications/ICTs in the Americas, under the principles of universality, equity, and affordability.

CITEL Strategic Plan 2018 – 2022: OBJECTIVES

1. To *increase telecommunication/ICT access, penetration, and coverage*.
2. To increase the *affordability* of telecommunication/ICT services and devices.
3. To *increase digital literacy and capabilities* relative to telecommunications/ICTs, as well as to build capacities to develop and keep communication networks in remote areas.
4. To *bridge the digital divide and reduce inequality*, particularly in underserved areas and regarding gender, disabilities, older persons, or persons with specific needs.
5. To *increase interoperability and interconnectivity*, of telecommunications/ICTs in the Americas, including international connectivity and harmonization of spectrum use.
6. To increase the budget of both the regular fund and specific fund, and the contributions of nonfinancial resources by the OAS.
7. To *increase the participation and positioning of CITEL* in the regional and international ICT/digital ecosystem.
8. To *increase the participation of Member States* and Associate Members in all meetings of CITEL.
9. To *improve CITEL's procedures and support tools* (website, search engine, data bank, access to hosted information, communication tools, etc.).





OAS | CITE

CITEL Structure

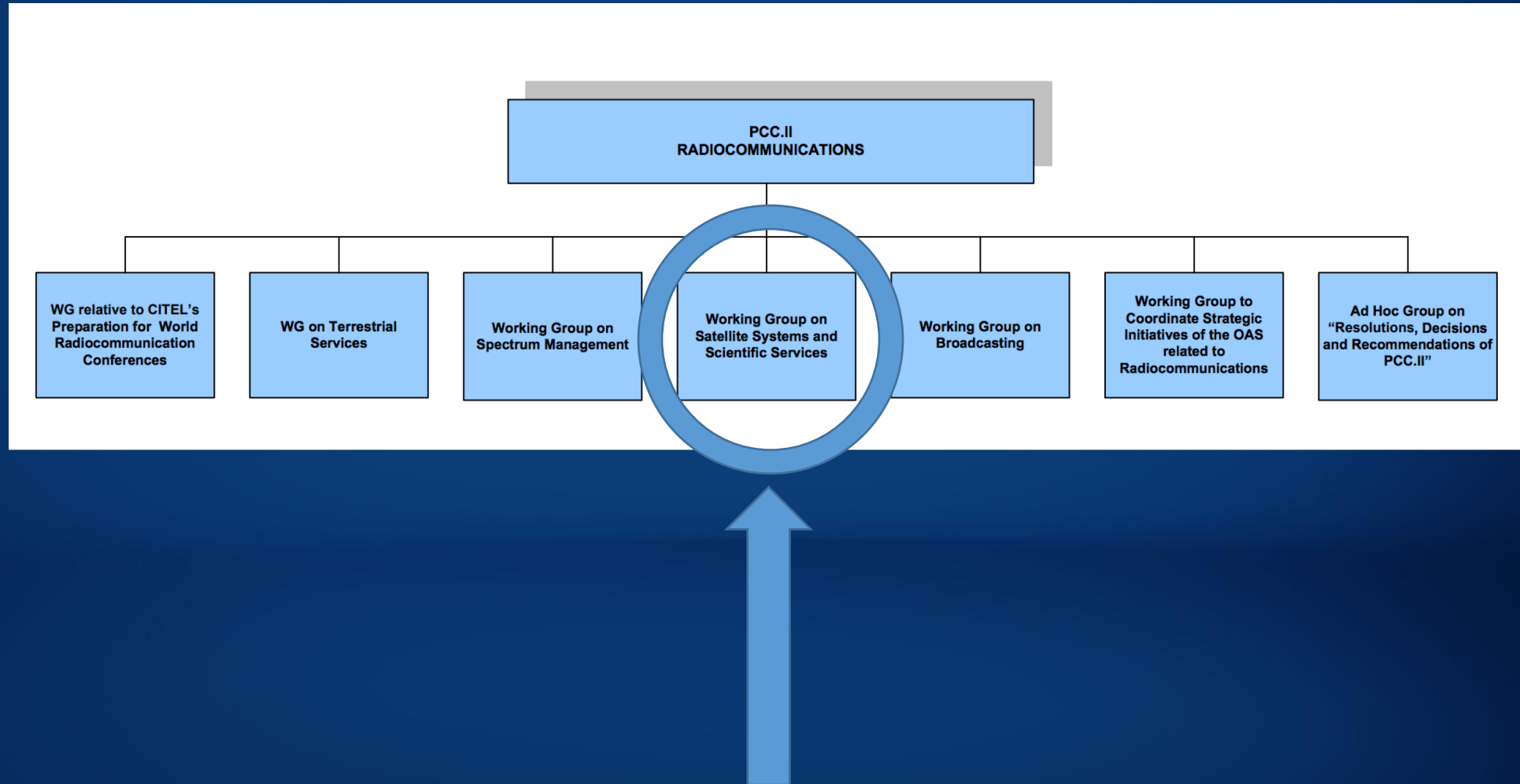
PCC.I - Telecommunications

PCC.II – Radiocommunications





OAS | CITEL



CITEL PCC.II Working Groups



OAS | CITEL

1. Working Group relative to CITEL's Preparation for World Radiocommunication Conferences.
2. Working Group on Terrestrial Services.
3. Working Group on Spectrum Management.
4. **Working Group on Satellite Systems and Scientific Services.**
5. Working Group on Broadcasting.
6. Working Group to Coordinate Strategic Initiatives of the OAS Related to Radiocommunications.
7. Ad Hoc Group on "Resolutions, Decisions and Recommendations of PCC.II"



Questions and Comments

esm@nsf.gov

Thank you!

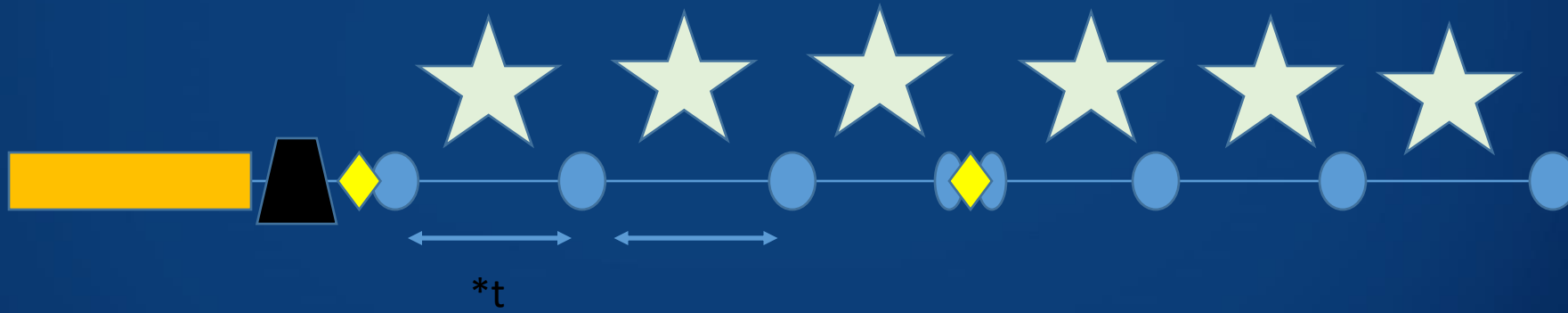


Credit: Sophia Dagnello, NRAO/AUI/NSF; NASA, STScI





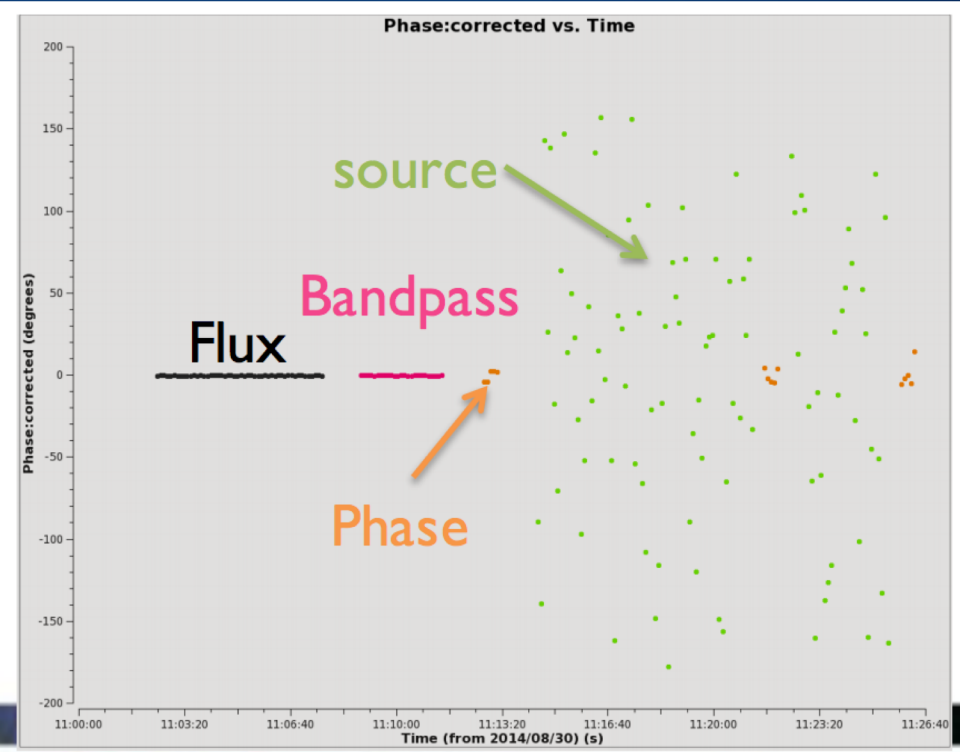
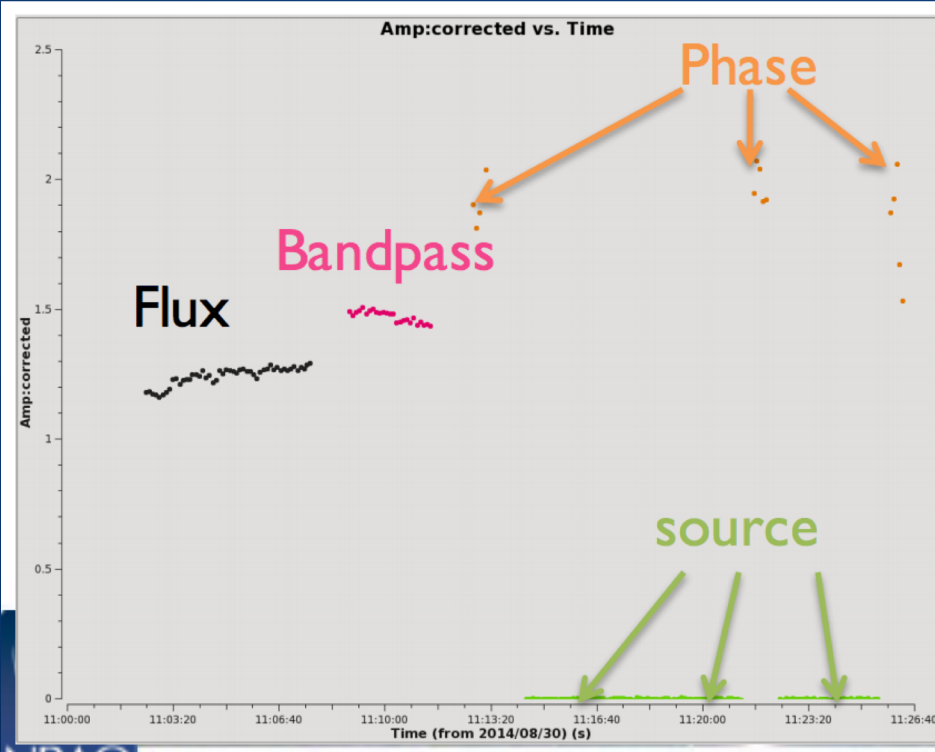
A sample Observation

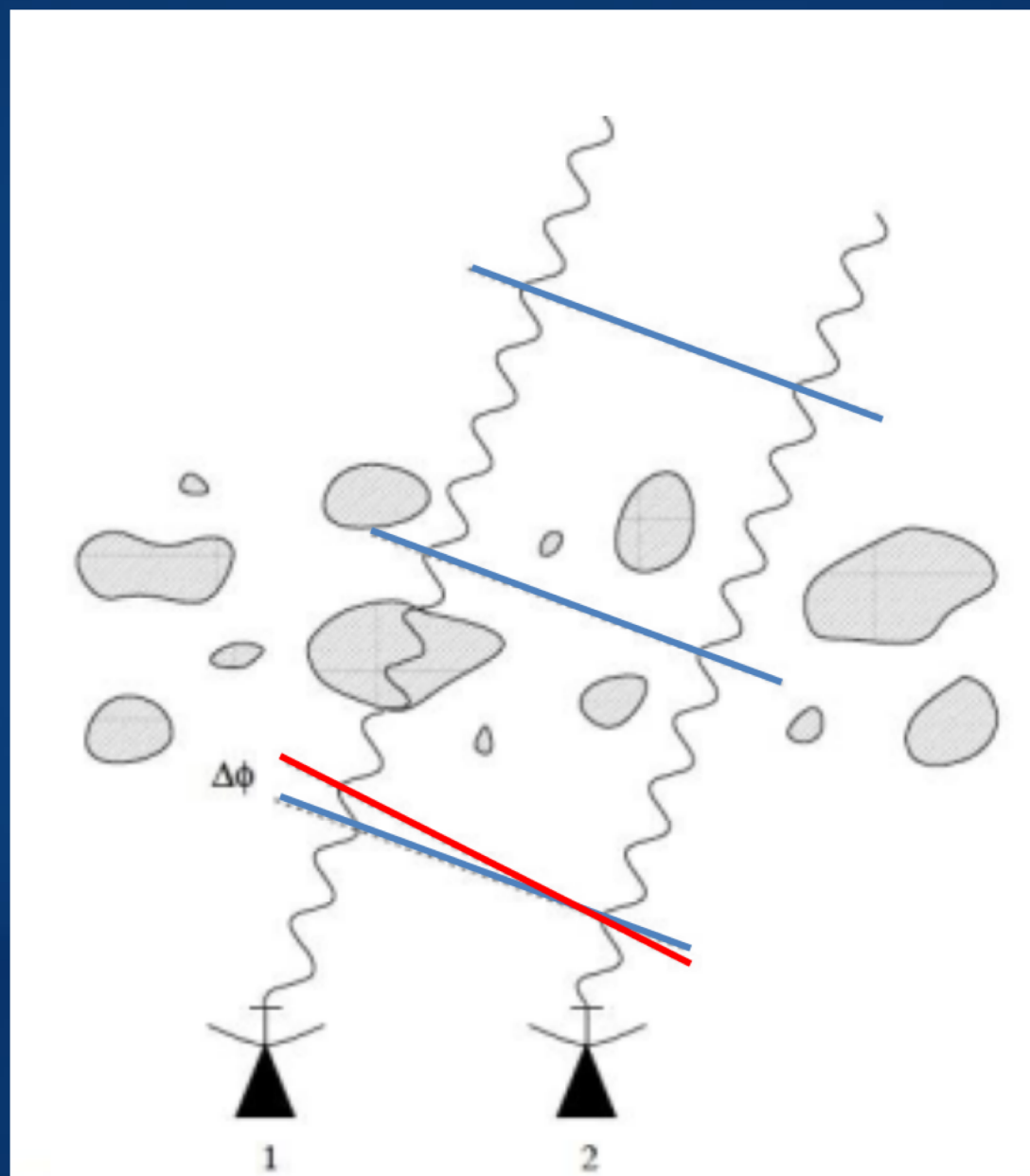


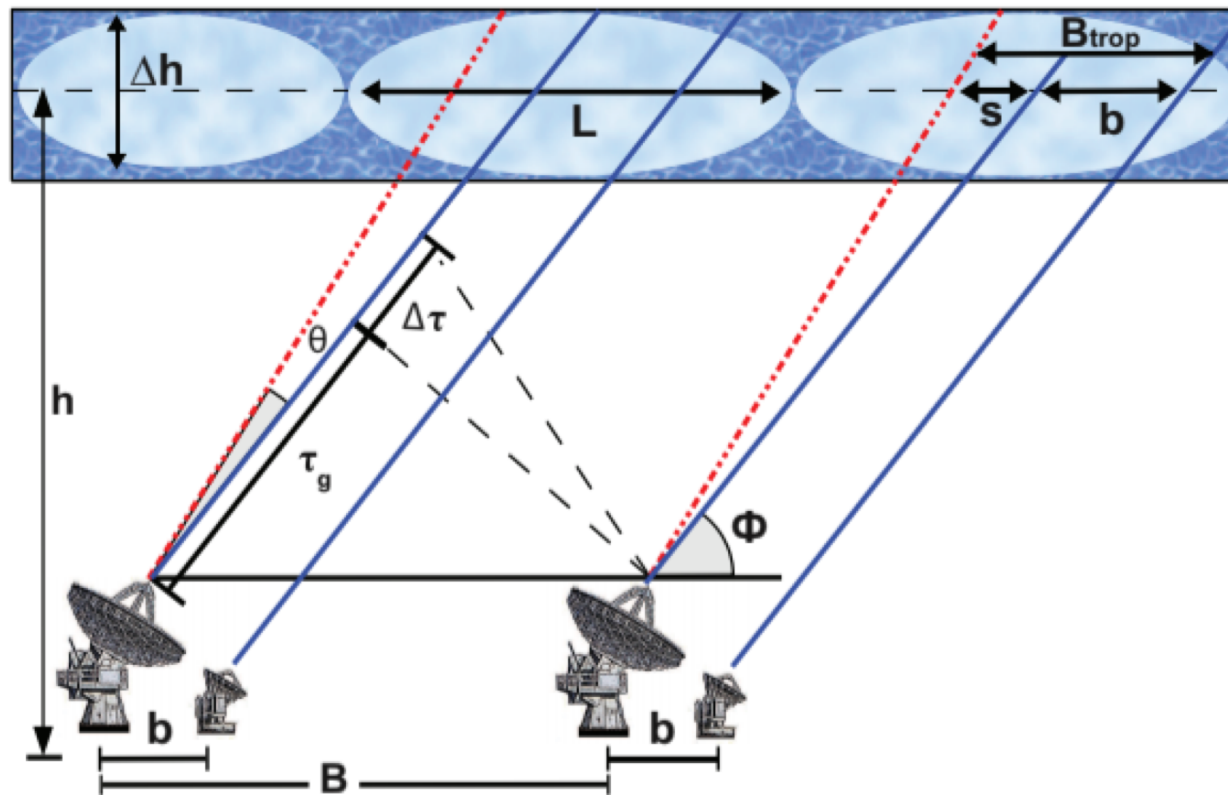
$*t$ - Time depends on...

- Frequency of observation
- Weather conditions
- Science goal
- Phase correction techniques (WVR, paired antennas, etc.)









Phase calibration

5

